

## 584.

ADDITION TO PROF. R. S. BALL'S PAPER, "NOTE ON A TRANSFORMATION OF LAGRANGE'S EQUATIONS OF MOTION IN GENERALISED COORDINATES, WHICH IS CONVENIENT IN PHYSICAL ASTRONOMY."

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THE formulæ may be established in a somewhat different way, as follows:—

Consider the masses  $M_1, M_2, \dots$

Let  $X_1, Y_1, Z_1$  be the coordinates (in reference to a fixed origin and axes) of the C.G. of  $M_1$ ;

$x_1, y_1, z_1$  the coordinates (in reference to a parallel set of axes through the C.G. of  $M_1$ ) of an element  $m_1$  of the mass  $M_1$ , and similarly for the masses  $M_2, \dots$ ; the coordinates  $(X_1, Y_1, Z_1), (X_2, Y_2, Z_2), \dots$  all belonging to the same origin and axes;

And let  $\dot{X}_1, \&c.$  denote the derived functions  $\frac{dX_1}{dt}, \&c.$

We have

$$T = S \frac{1}{2} m_1 [(\dot{X}_1 + \dot{x}_1)^2 + (\dot{Y}_1 + \dot{y}_1)^2 + (\dot{Z}_1 + \dot{z}_1)^2] \\ + S \frac{1}{2} m_2 [(\dot{X}_2 + \dot{x}_2)^2 + (\dot{Y}_2 + \dot{y}_2)^2 + (\dot{Z}_2 + \dot{z}_2)^2] \\ :$$

or since  $S m_1 x_1 = 0, \&c.$ , and therefore also  $S m_1 \dot{x}_1 = 0, \&c.$ , this is

$$T = \frac{1}{2} M_1 (\dot{X}_1^2 + \dot{Y}_1^2 + \dot{Z}_1^2) + S \frac{1}{2} m_1 (\dot{x}_1^2 + \dot{y}_1^2 + \dot{z}_1^2) \\ + \frac{1}{2} M_2 (\dot{X}_2^2 + \dot{Y}_2^2 + \dot{Z}_2^2) + S \frac{1}{2} m_2 (\dot{x}_2^2 + \dot{y}_2^2 + \dot{z}_2^2) \\ :$$



Write  $u, v, w$  for the coordinates of the c.g. of the whole system: then

$$M_1 X_1 + M_2 X_2 + \dots = (M_1 + M_2 \dots) u,$$

$$M_1 Y_1 + M_2 Y_2 + \dots = (M_1 + M_2 \dots) v,$$

$$M_1 Z_1 + M_2 Z_2 + \dots = (M_1 + M_2 \dots) w;$$

and thence

$$M_1 \dot{X}_1 + M_2 \dot{X}_2 + \dots = (M_1 + M_2 \dots) \dot{u},$$

$$M_1 \dot{Y}_1 + M_2 \dot{Y}_2 + \dots = (M_1 + M_2 \dots) \dot{v},$$

$$M_1 \dot{Z}_1 + M_2 \dot{Z}_2 + \dots = (M_1 + M_2 \dots) \dot{w};$$

and thence

$$\begin{aligned} T - \frac{1}{2} (M_1 + M_2 + \dots) (\dot{u}^2 + \dot{v}^2 + \dot{w}^2) \\ = \frac{1}{M_1 + M_2 \dots} \{ M_1 M_2 [(\dot{X}_1 - \dot{X}_2)^2 + (\dot{Y}_1 - \dot{Y}_2)^2 + (\dot{Z}_1 - \dot{Z}_2)^2] \\ : \\ + S \frac{1}{2} m_1 (\dot{x}_1^2 + \dot{y}_1^2 + \dot{z}_1^2) \\ + S \frac{1}{2} m_2 (\dot{x}_2^2 + \dot{y}_2^2 + \dot{z}_2^2) \\ : \end{aligned}$$

or, representing the function on the right-hand side by  $T'$ , this is

$$T = \frac{1}{2} (M_1 + M_2 + \dots) (\dot{u}_2 + \dot{v}_2 + \dot{w}_2) + T' \dots = T_0 + T'.$$

Suppose the positions are determined by means of the  $6n$  coordinates  $((q))$ ; the equations of motion are each of them of the form

$$\frac{d}{dt} \cdot \frac{dT_0}{dq} - \frac{dT_0}{dq} + \frac{d}{dt} \cdot \frac{dT'}{dq} - \frac{dT'}{dq} = - \frac{dV}{dq}.$$

But these admit of further reduction; the part in  $T_0$  depends upon three terms, such as

$$\frac{d}{dt} \left( \dot{u} \frac{d\dot{u}}{dq} \right) - \dot{u} \frac{d\dot{u}}{dq}, = \frac{d\dot{u}}{dt} \frac{d\dot{u}}{dq} + \dot{u} \left( \frac{d}{dt} \frac{d\dot{u}}{dq} - \frac{d\dot{u}}{dq} \right).$$

But we have  $u$  a function of  $((q))$ , and thence

$$\frac{d\dot{u}}{dq} = \frac{du}{dq}, \text{ or } \frac{d}{dt} \frac{d\dot{u}}{dq} - \frac{d\dot{u}}{dq}, = \frac{d}{dt} \frac{du}{dq} - \frac{du}{dq}, = 0,$$

or the term is simply

$$= \frac{d\dot{u}}{dt} \frac{d\dot{u}}{dq}.$$

The equation thus becomes

$$(M_1 + M_2 \dots) \left( \frac{d\dot{u}}{dt} \frac{d\dot{u}}{dq} + \frac{d\dot{v}}{dt} \frac{d\dot{v}}{dq} + \frac{d\dot{w}}{dt} \frac{d\dot{w}}{dq} \right) + \frac{d}{dt} \frac{dT'}{dq} - \frac{dT'}{dq} = - \frac{dV}{dq}.$$



Suppose now that  $T'$ ,  $V$  are functions of  $6n - 3$  out of the  $6n$  coordinates ( $(q)$ ), and of the differential coefficients  $\dot{q}$  of the same  $6n - 3$  coordinates, but are independent of the remaining three coordinates and of their differential coefficients; then, first, if  $q$  denotes any one of the three coordinates, the equation becomes

$$\frac{d\dot{u}}{dt} \frac{d\dot{u}}{d\dot{q}} + \frac{d\dot{v}}{dt} \frac{d\dot{v}}{d\dot{q}} + \frac{d\dot{w}}{dt} \frac{d\dot{w}}{d\dot{q}} = 0;$$

or, better,

$$\frac{d\dot{u}}{dt} \frac{d\dot{u}}{d\dot{q}} + \frac{d\dot{v}}{dt} \frac{d\dot{v}}{d\dot{q}} + \frac{d\dot{w}}{dt} \frac{d\dot{w}}{d\dot{q}} = 0;$$

and the three equations of this form give

$$\frac{d\dot{u}}{dt} = 0, \quad \frac{d\dot{v}}{dt} = 0, \quad \frac{d\dot{w}}{dt} = 0,$$

viz. these are the equations for the conservation of the motion of the centre of gravity.

And this being so, then, if  $q$  now denotes any one of the  $6n - 3$  coordinates, each of the remaining equations assumes the form

$$\frac{d}{dt} \cdot \frac{dT'}{d\dot{q}} - \frac{dT'}{dq} = - \frac{dV}{dq},$$

viz. we have thus  $6n - 3$  equations for the relative motion of the bodies of the system.